

Heating Rates in 3D Cloud Fields

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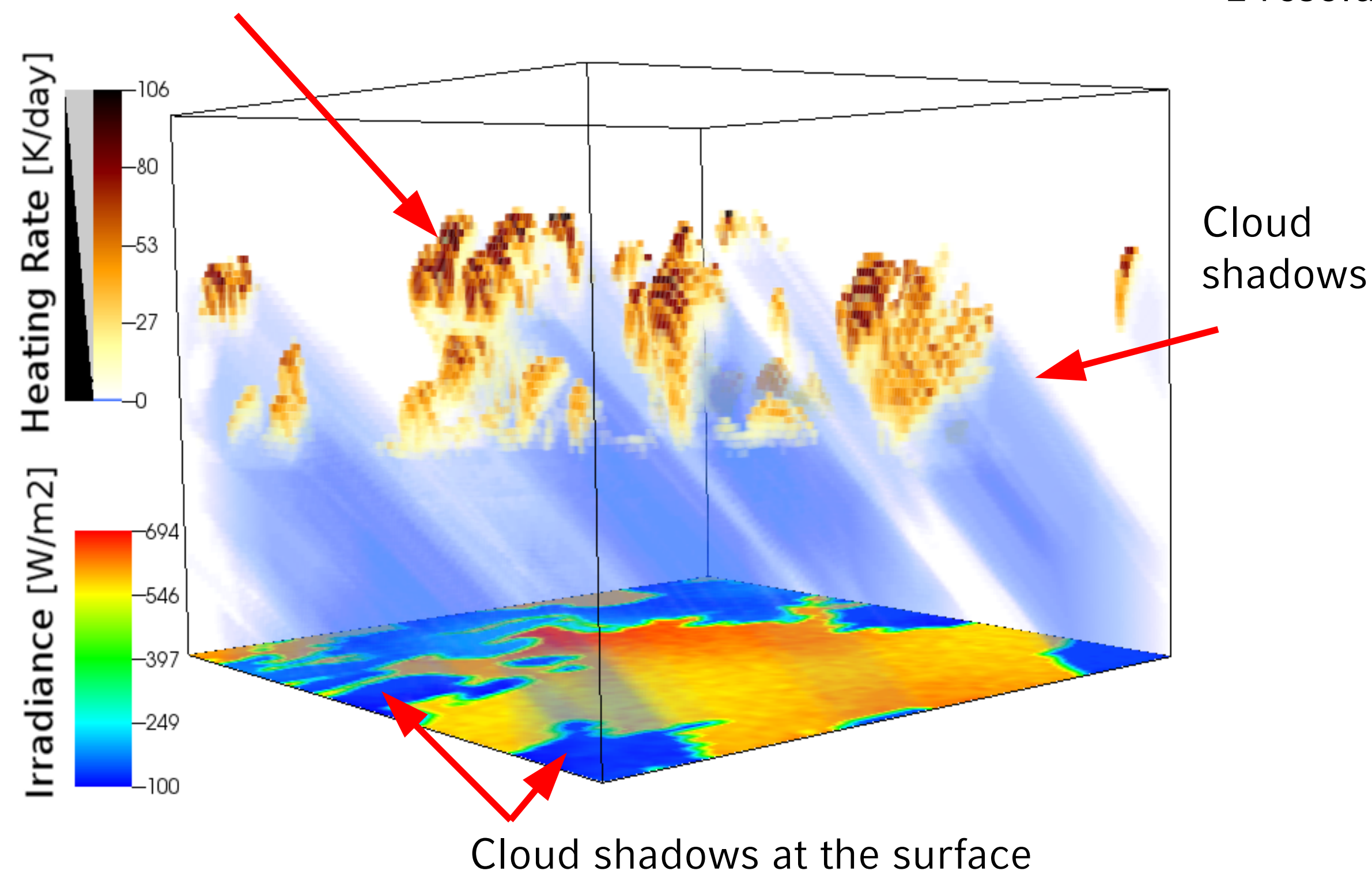
Carolin Klinger

3D Solar Heating Rates

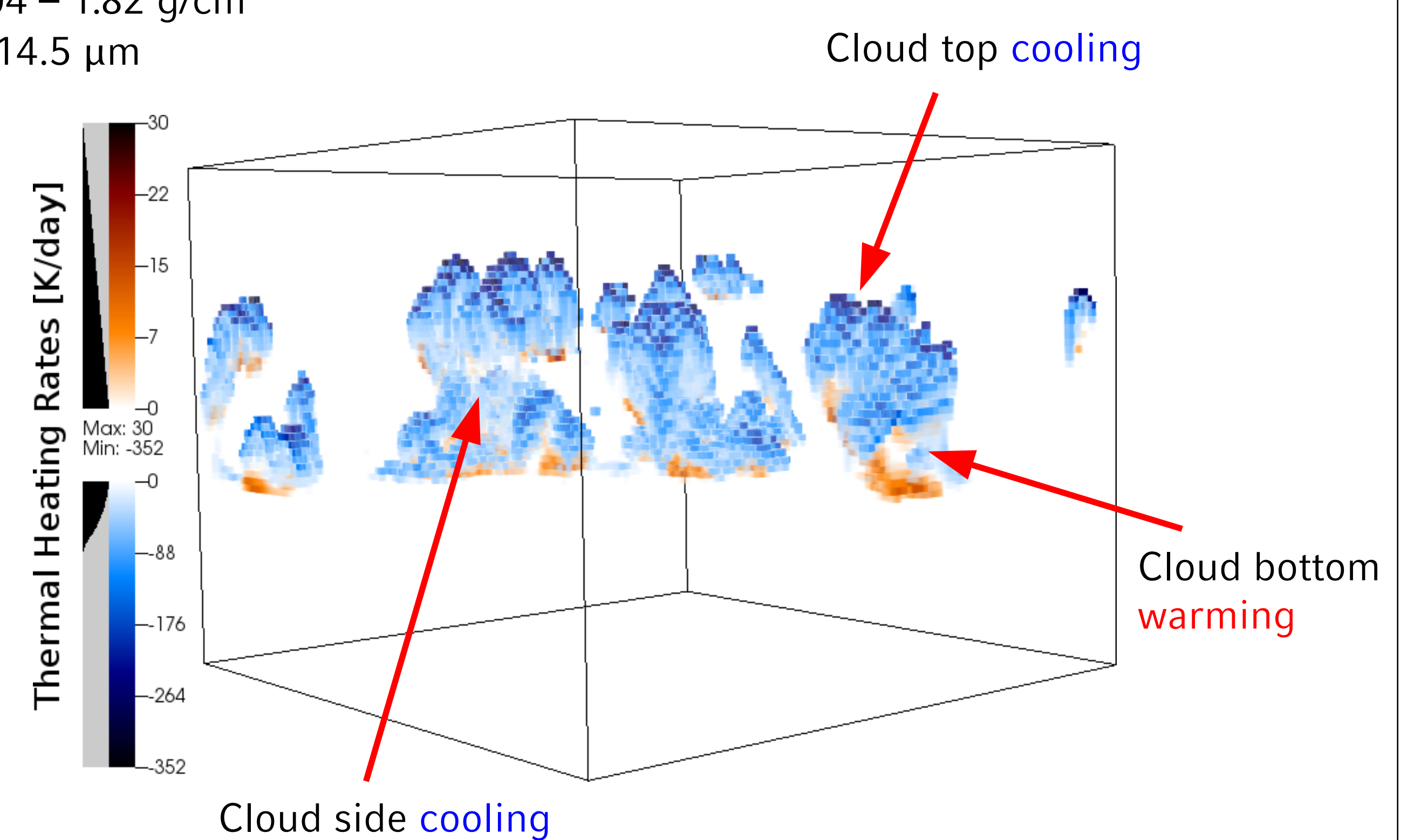
I3RC Cumulus Cloud Field:

96 * 96 pixels
 x/y resolution: 67 m
 z resolution: 40 m
 Cloud Height: 1.04 - 2.4 km
 LWC: 0.004 - 1.82 g/cm³
 R_{eff}: 4 - 14.5 μm

Warming at illuminated cloud side/top



3D Thermal Heating Rates



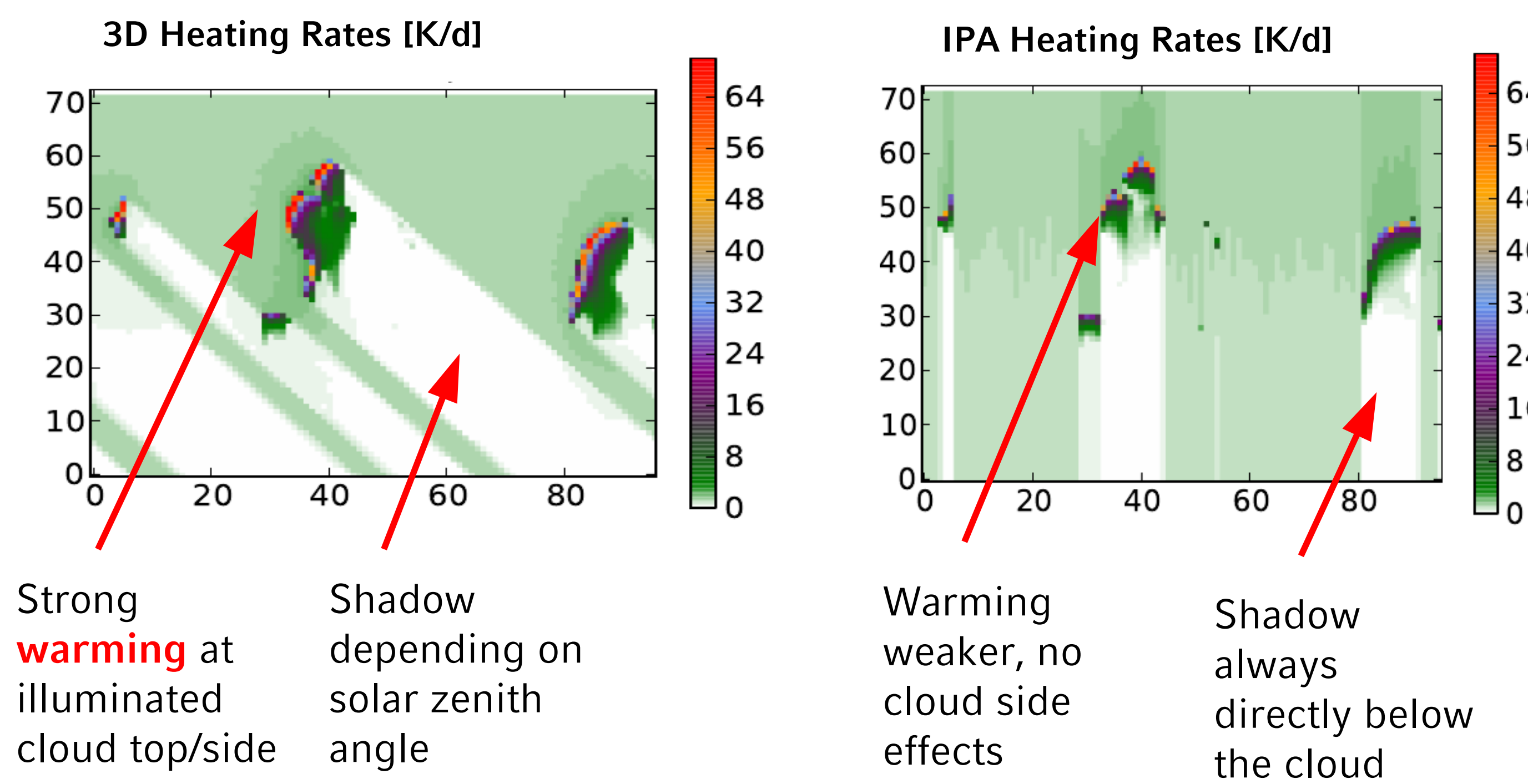
- Clouds reduce the amount of radiation that reaches the surface
- Heat flux from surface to the atmosphere is reduced
- Absorption of solar radiation at illuminated cloud tops/cloud sides → warming effect

- Influence on:
- Cloud microphysics
 - Dynamics
 - Cloud development

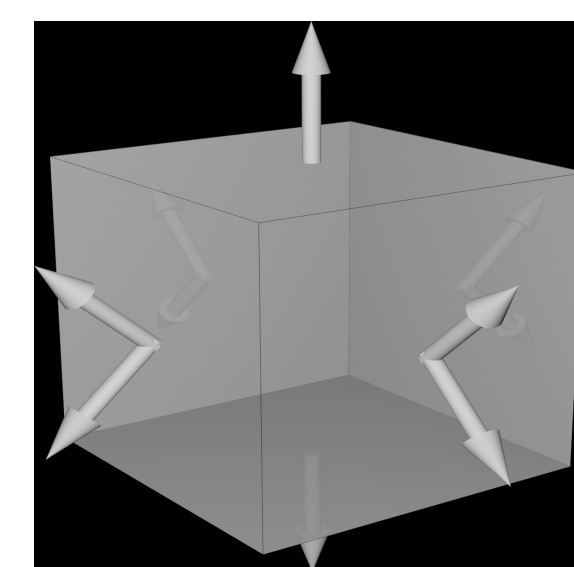
- Main interactions: thermal absorption and emission of the clouds and the atmosphere
- Cloud side and cloud top cooling
- Cloud bottom warming

Independent Pixel Approximation

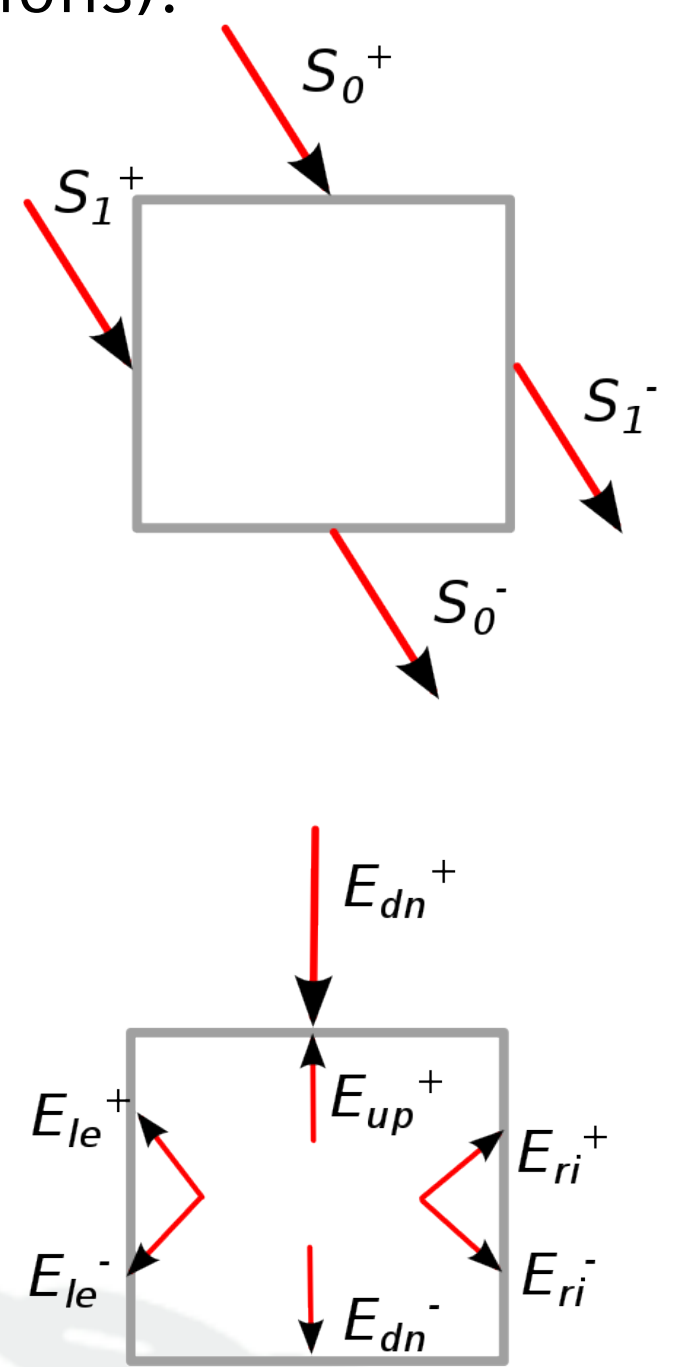
Comparison of 3D calculation and the commonly employed independent pixel approximation (IPA) in the solar spectral range for the I3RC cumulus cloud field.



Concept for Fast 3D-RT Solver

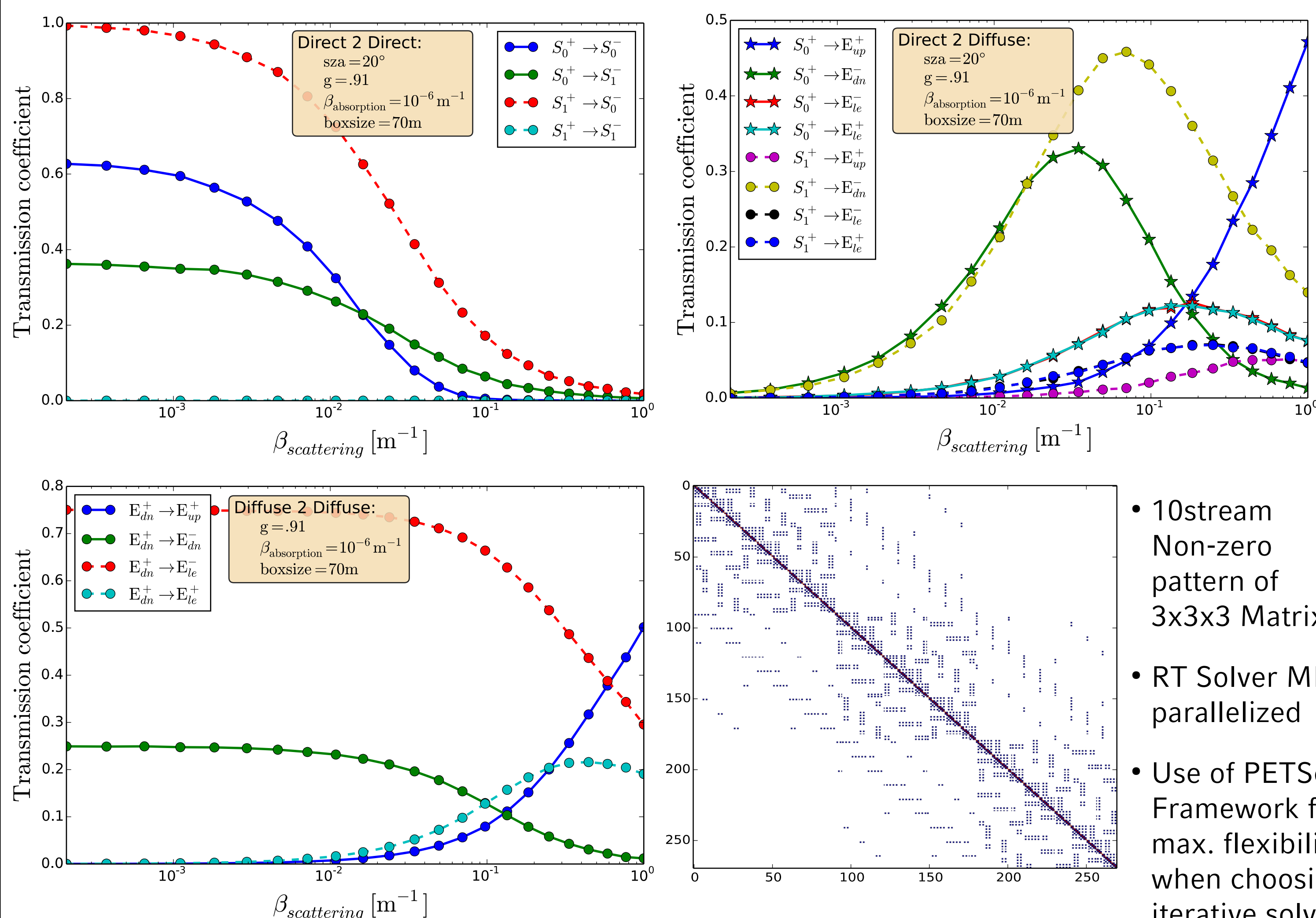


- Generalization of the twostream method to 3 dimensions
- Minimum required streams for 3D-RT(directions):
 - 3 streams for direct radiation ($S_{0,1,2}$)
 - 10 streams for diffuse radiation ($E_{dn,up,...}$)



- Obtain transport coefficients for arbitrary geometries and optical properties with exact Monte Carlo methods and store in LookUpTable
- Degrees of freedom of transport coefficients for one box are aspect ratio, optical properties (absorption / scattering coefficients and asymmetry parameter) and sun angles (zenith and azimuth)
- Couple streams in linear equation system and solve iteratively

Transport Coefficients for Single Box

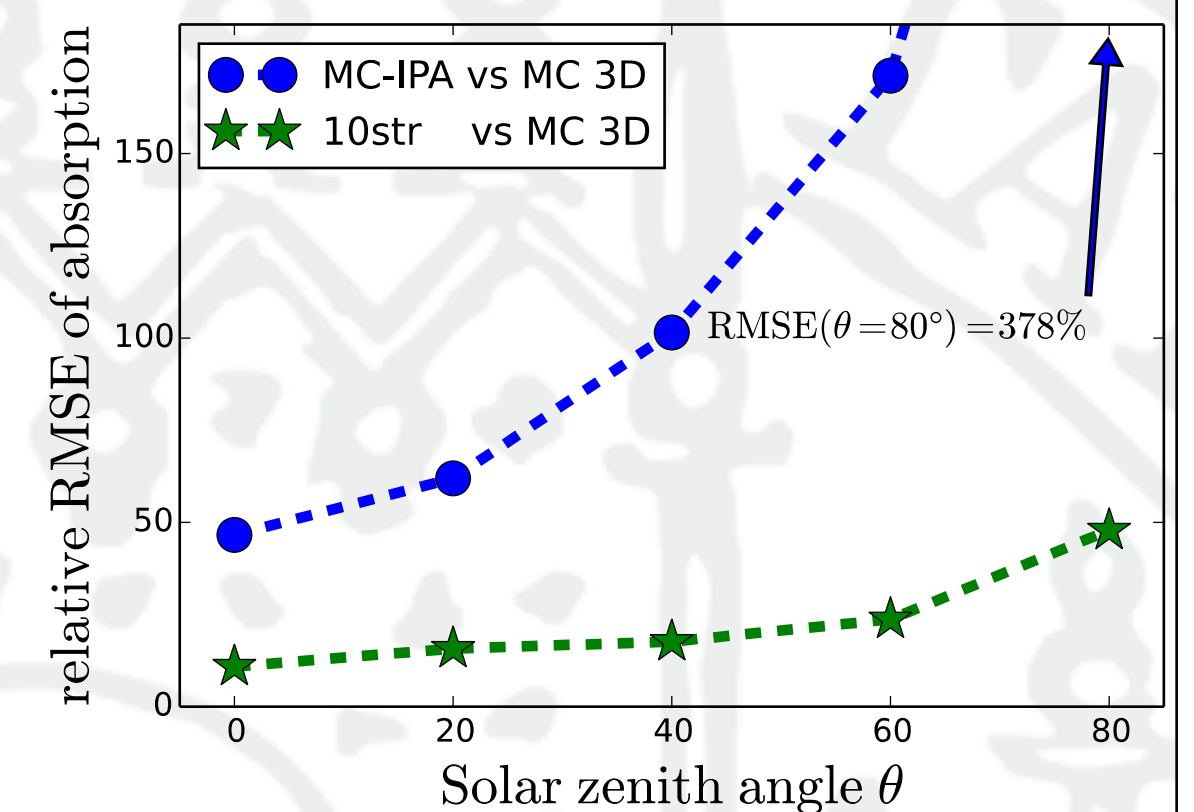


- 10stream Non-zero pattern of 3x3 Matrix
- RT Solver MPI-parallelized
- Use of PETSc Framework for max. flexibility when choosing iterative solver

Performance of New Solver

Can it calculate 3D absorption?

- Root-Mean-Square-Error shows significant improvement (here shown is the RMSE for I3RC scene from above)
- Decrease in precision for high solar zenith angles due to numerical diffusion in direct part



What are the computational costs?

- Calculations done on workstation with 8 cores
- Dashed lines (zero-guess) show an approximative increase of total runtime by a factor of 12
- If started with a good initial guess (e.g. last time step +20%noise perturbation), computation may be significantly faster

